## Average distance between sun and earth smoon and earth

| 1 | Average Distance between sun and earth | $\begin{gathered} 14,95,98,262 \\ \text { KMs. } \end{gathered}$ | 15 Crores KMs Appx. | $\begin{aligned} & \text { For more details } \\ & \text { see Page - } 4 \text { (1) } \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: |
| 2 | Average Distance between Moon and Earth | 3,84,403 KMs. | 3.85 Lakhs KMs. <br> (Appx.) | Sun is 390 times more far (3.85L Kms X 390=15 Cr.) |

3 Earth's average orbital speed is about 30 KMs per second. for $1,10,000 \mathrm{KMs}$. per hour\}, whereas rotation Speed is $1670 \mathrm{~K} . \mathrm{Ms}$. per Hour \{27.8 KMs per Minute, 0.46 Kms. Per Sec.Distance of earth travel in one year (Length of orbit) 94 Cr . KMs

## Circumference and Diameter of Earth, Moon, Sun

1 Circumstances of the earth at 40075 KMs For more details equator(The distance around the Earth at the Equator)
2 Diameter of the earth
3 Circumference of the Moon Diameter of Moon

$$
12.756 \mathrm{kMs}
$$

Circumference of the Sun 10961 KMs 3,474.2 KM
s
43.79 Lakhs кмs
13.91 Lakhs

кмs

| $\mathbf{1}$ | Time takes for one rotation for Earth, Moon and Sun <br> Rotation-A day | 23 Hrs. 56 Mts. 04.09053 Secs. |
| :--- | :--- | :--- |
| $\mathbf{2}$ | Time takes for moon for one <br> Rotation- A day | 27.322 days (Sidereal day) |
| $\mathbf{3}$ | Time takes for Sun for one <br> Rotation- A Day | 25.38 days |
|  | Astronomers measure the rotation rate of the Sun from an arbitrary position <br> of $26^{\circ}$ from the equator; around the point where sunspots are observed. At <br> this point, it takes 25.38 days to rotate and return to the same spot in <br> space. So that would beone DAY on the Sun for all practical purposes. |  |

## Average moving speed of Earth, Moon and Sun (Rotation on its axis)

| 1 | Average moving speed of earth f (Rotates on its axis) (At the equator) \}. | 1670 K.Ms. pe per Minute, 0.4 Circumstances 40075 KMs/2 | ur $\{27.8 \mathrm{KMs}$ ns. Per Sec. th at equator is ay=1670 KMs. | How fast are you moving when sitting still |
| :---: | :---: | :---: | :---: | :---: |
| 2 | Average moving speed of moon \{ (Rotates on its axis) (Appx.)\} | 24 KMs per Hour <br> \{The equatorial circumference of the Moon is 10961 KMs./456 Hours (27 Days 7 H 43.1 M 2.9 Secs) (Appx. 456 Hours) \}= 24 KMs per Hour. (For more details about moon See page No. 8\} |  |  |
| 3 | Average moving speed of SUN (Rotates on its axis) | On average, the sun rotates on its axis once every 27 days. \{"Since the sun is a ball of gas/plasma(Sun is n't a solid object like a planet), it does not have to rotate rigidly like the solid planets and moons do," \{Its rotation is harder to pinpoint |  |  |
| Orbital Period (To take a round) of earth, moon and Sun |  |  |  |  |
| 1 | Orbital Period of earth (To take a <br> round around Sun)$\quad$365.25 days <br> For more details see Page - 4 (6) |  |  |  |
| 2 | Orbital Period o | moon (To take a | 27.321582 Earth days |  |


|  | round around earth) \{27 1/3 days (Appx.]) \} |  | $\begin{gathered} 2 \text { Days } 7 \text { H } 43.1 \text { M } 2 \\ \text { Secs. } \end{gathered}$ |
| :---: | :---: | :---: | :---: |
|  | Moon takes 27.32 days to rotate on its axis and orbital period of moon is same time to rotate on its axis, as a result the moon always presents it's same side to the earth. Scientists call this synchronous rotation. What this means is that the Moon is tidally -locked with Earth. |  |  |
| 3 | L PERIOD OF SUN take around s the Sun move und the Milky <br> ?) (Answer is Yes) \} | Our whole sol the Milky Way We are moving km/hr. But eve about 230 mil | orbits around the center of or more details see Page - 6) verage velocity of 828,000 high rate, it still takes us to make one complete orbit he Milky Way. |
| On Page 4 ( 8 \& 9): Information about Equinox (Equal day and night) and Solstice (Long day and Long night are available |  |  |  |

UNIVERSE - INTERESTING FEATURES (Cntd.)

|  | How many <br> minutes does the <br> Sun light take to reach the E arth $=8$ Mts. 20 Secs. | Average Distance between sun and earth is14,95,98,262 KMs. /(Divided by) <br> Travelling speed of Light is 2.99 Lakh KMs per Sec. $=500 \mathrm{KMs}$ per Sec. $=$ $\begin{array}{r} 500 / 60=8.33 \text { Mts. } \\ \text { Mts. } 20 \text { Secs.) } \end{array}$ | Another theory is time taken for the journey (from the Sun's surface to the Earth's) will vary between 8.14 and 8.42 minutes, |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 3 | Travelling speed of Light (TSL) |  | SPEED OF Mtrs . <br>  P/Sec |  |  |
|  |  | 2,99,792.458(3 Lakhs. KMs per/Second) | Light | 30 Cr . |  |
|  | B. Light minute (Light Second X 60) | 1.79,87,547.48 | Explosio <br> n | 8000 |  |
|  | C. 1 Light hour (Light Minute X 60) | 107, 92, 52 , 848.8 | Bullet | 810 |  |
|  |  | $2,590,20,68,371.19$ | Sound | 343 |  |
|  | $\frac{\text { E. } 1 \text { Light-year }}{\substack{\text { Light day } X \\ 365)}}$ | $\begin{gathered} \mathbf{9 , 4 6 , 0 7 3 , 0 4 , 7 2 , 5 8 0 . 7 9} \\ \text { (9.5 Lakh Crores or } \\ \text { say } 10 \text { Trillion) } \end{gathered}$ | nommanammana |  |  |
| 4 | General <br> Maximum speed of Passenger flight | 900 to 1000 KMs per hour \{If way say it is 1000 KMs per Hour. (Per Minute is 1000/60 = 16.66 KM .) <br> Per Secondit is <br> $16.66 / 60=0.277 \mathrm{Kms}$. Per Second Meters is 0.277 KMsX 1000 Mtrs.) 277 Mtrs./Sec\} |  | $\begin{gathered} 277 \\ \text { Mtrs./Sec } \end{gathered}$ |  |
|  | Cruising altitude of Flight (Vertical distance from Sea level) | 25000 Ft. to 4000 Ft. depend upon type and size of Aircraft <br> Generally - say it goes up to the Ht. of 10 KMs. (32808 Ft.) <br> (Concorde flight travel at the Ht. of 60000 Ft.) <br> Ht. of Mount Everest is 60000 Ft (8288 Mts.=8.2 Mts) |  |  |  |




|  | INTERESTING FEATURES ABOUT EARTH |  |
| :---: | :---: | :---: |
| 1 | Average Distance between sun and earth | 14,95,98,262 KMs. |
|  | The distance from the Earth to the Sun varies because the Earth's orbit around the Sun is elliptical. So, the time will vary by about 3\% depending on the Earth's distance from the Sun |  |
| 2 | Average Distance between Moon to Eart | 3,84,403 KMs |
| 3 | The distance around the Earth at the Equator (It is circumference) | 40,075 KMs. (40K) |
|  | The equatorial circumference of the Earth is $40,075 \mathrm{~km}$. This is the distance around the equator of the Earth. If you measure the circumference of the Earth, while passing through the poles, the distance is only 40,007 km. This is because the Earth isn't a perfect sphere. It's rotating rapidly, which causes the equator to bulge out. |  |
| 4 | ameter of the ea | 12.756 KMs |
|  | The equatorial diameter of the Earth is $12,756 \mathrm{~km}$. This is the diameter of the Earth measured from one side of the Earth, passing through the center. If you go from pole to pole through the center, the distance is only 12,713 kms. |  |
| 3 | ```Average moving speed of earth {(Rotates on its axis) (At the equator)}.``` | 1670 K.Ms. per Hour |
|  | Earth's spin is constant, but the speed depends on what latitude you are located at. Here's an example. The circumference (At the equator)) is roughly 40,070 kilometers, according to NASA. If you estimate that a day is 24 hours long, you divide the circumference by the length of the day. This produces a speed at the equator of about1,670 km/h. |  |
| 4 | Orbital Period of earth (To take a round around Sun) | 365.25 days |
| 5 | Time for taking One Rotation on Axi | $\begin{gathered} 23 \text { Hrs. } 56 \text { Mts. } 04.09053 \\ \text { Secs. } \end{gathered}$ |
| 6 | Time for taking One Revolution around Sun: | 365.25 days |
|  | - How long does the Earth take to travel around the Sun? Obviously, the answer is one year or 365.25 days. It is not so simple though as there are a number of definitions of a year. For example, <br> - Tropical year, which is from equinox to equinox, that is from the time the Sun crosses the celestial equator from south to north to the next time 365.24219 days <br> - Sidereal year, from one time a particular star is in a given position to the next time 365.25636 days <br> - Anomalistic year, from the time the Earth is at its closest to the Sun to the next time 365.25964 days |  |
| 7 | Water vs. Land: 70.8\% Water, 29.2\% Land |  |
| A | Surface Area of Total Earth (Land + Water) | $\begin{gathered} \hline 51,01,00,500 \text { Sq. Kms. }(51 \\ \text { Cr. }) \end{gathered}$ |
| B | Surface Area of Earth (Only Land) | $\begin{gathered} 14,88,51,000 \mathrm{Sq} . \\ \mathrm{Kms.}(14.9 \mathrm{Cr}) \end{gathered}$ |
| 8 | Equinox (All most equal day and eq twice each year: around 20 Mar <br> - The equinoxes are the only times when to the equator. | l night )- This occurs and 23 September. <br> he Sun is perpendicular |



|  | dioxide and water |
| :---: | :---: |
|  | Chemical Composition of the Earth: 34.6\% Iron, 29.5\% Oxygen, 15.2\% Silicon, 12.7\% Magnesium, 2.4\% Nickel, 1.9\% Sulphur, and 0.05\% Titanium. |
|  | Distance around the Earth at the Equator (It is circumference) is $40,075 \mathrm{KMs}$. <br> Therefore Minimum distance between any two places on the globe should not be more than around. : 20,000 KMs. <br> The Shanghai city of China of Asian continent (Located in north east place) and Buenos Aries city of Argentina city (Located in South-West corner) of South America continent (earth) distance between two cities is 19,644 KMs. |

## StarChild Question of the Month for February 2000

## Question:

Does the Sun move around the Milky Way?
$\qquad$

## Answer:

Yes, the Sun - in fact, our whole solar system - orbits around the center of the Milky Way Galaxy. We are moving at an average velocity of $828,000 \mathrm{~km} / \mathrm{hr}$. But even at that high rate, it still takes us about 230 million years to make one complete orbit around the Milky Way!

The Milky Way is a spiral galaxy. We believe that it consists of a central bulge, 4 major arms, and several shorter arm segments. The Sun (and, of course, the rest of our solar system) is located near the Orion arm, between two major arms (Perseus and Sagittarius). The diameter of the Milky Way is about 100,000 light-years and the Sun is located about 28,000 light-years from the Galactic Center. You can see a drawing of the Milky Way below which shows what our Galaxy would look like "face-on" and the direction in which it would spin as viewed from that vantage point. Also shown, is the location of the Sun in the big picture view of our Galaxy.


It is interesting to note that recent observations by astronomers suggest that the Milky Way is in fact a "barred spiral galaxy", not just a "spiral galaxy". This means that rather than a simple spherical bulge of gas and stars at its center, it has instead a "bar of stars" crossing the central bulge. It might look something like the image shown below of the barred spiral galaxy known as NGC1073. But we still rotate around the center just the same!

```
    The Milky Way (Pala Puntha) is a barred spiral galaxy with a
diameter between 150,000 and 200,000 light-years (ly).[22][23][24][25] It
    is estimated to contain 100-400 billion stars[26][27] and more than
    100 billion planets.[28][29] The Solar System is located at a radius
        of \(26,490( \pm 100)\) light-years from the Galactic Center,
        A galaxy ( నక్షత్రవిధి
            ) is a gravitationally bound system of stars, stellar
        remnants, interstellar gas, dust, and dark matter.[1][2] The word
        galaxy is derived from the Greek galaxias (voג \(\quad\) (í \(\alpha\) ) , literally
    "milky", a reference to the Milky Way. Galaxies range in size
        from dwarfs with just a few hundred million (10 \({ }^{8}\) ) stars
    to giants with one hundred trillion (1014) stars, [3] each orbiting
        its galaxy's center of mass.
```

ON AVERAGE, THE SUN ROTATES ON ITS AXIS ONCE EVERY 27 DAYS.
It takes 24 hours for the Earth to make a full rotation, but since the sun isn't a solid object like a planet, its rotation is harder to pinpoint.
"Since the sun is a ball of gas/plasma, it does not have to rotate rigidly like the solid planets and moons do," according to NASA.

In fact, our gaseous sun is divided into different zones and layers, with each of our host star's regions moving at varying speeds. On average, the sun rotates on its axis once every 27 days. However, its equator spins the fastest and takes about 24 days to rotate, while the poles take more than 30 days. The inner parts of the sun also spin faster than the outer layers, according to NASA.

## SUN



## The Average Motion of the Moon

The Moon moves around the Earth in an approximately circular orbit, going once around us in approximately 27.3 days, or one sidereal period of revolution. As it does this its position changes, relative to the stars.
Since there are 360 degrees in a circle, the Moon moves (on the average) 360 / 27.3 or 13.2 degrees per day relative to the stars, which is just over half a degree per hour, and approximately equal to its apparent size. This means that from night to night the Moon moves a little more than one hand-width to the East (the direction of its motion around the Earth) relative to the stars, and from hour to hour it moves about one diameter to the East, among the stars.


An approximate representation of the motion of the Moon around the Earth. Moving once around in 27.3 days, its average movement is about 13.2 degrees per day, or 92 degrees per week. (As is usual in such diagrams the sizes of the Earth and Moon are exagerrated, in comparison to their separation.)

## LIGHT YEAR

Light year (symbol: ly), is a unit of length equal to just under 10 trillion kilometres (or about 6 trillion miles). As defined by the International Astronomical Union(IAU), a light-year is the distance that light travels in a vacuum in one Julian year. ${ }^{[1]}$ The light-year is mostly used to measure distances to stars and other distances on a galacticscale, especially in non-specialist and popular science publications. The preferred unit inastrometry is the parsec (approximately 3.26 light-years), because it can be more easily derived from, and compared with, observational data. ${ }^{[11\}}$ Note that the light-year is a measure of distance (rather than, as is sometimes misunderstood, a measure of time). $\}$

1 light-year
$=9460730472580800$ metres $($ exactly $) \approx 5878625$ million miles $\approx 63241.1$ astronomical units $\approx 0.306601$ parsecs. The figures above are based on a Julian year (notGregorian year) of exactly 365.25 days (each of exactly 86400 SI seconds, totalling 31557600 seconds) ${ }^{[2]}$ and a defined speed of light of $299792458 \mathrm{~m} / \mathrm{s}$, both included in the IAU (1976) System of Astronomical Constants, used since 1984. ${ }^{[3}$
6.Gravity is the force of attraction that exists between any two objects.

- the more massive two objects are, the stronger the force of gravity between them
- the farther apart two objects are, the weaker the force of gravity between them
- If an object is launched from the surface of the Earth, it needs to reach a certain speed called the escape velocity in order to break free of the Earth's gravity. This speed is about 7 miles per second, or 25,000 miles per hour. If the object doesn't reach escape velocity, it will either crash back into the Earth, or enter into orbit around it, as satellites or the space shuttle do.
- To completely escape the earth's gravitational pull the crew must reach a distance of at least 400 km away, that is why the iss (international space station) is at an average distance of 250 km from the earth constantly. The pull is strong enough to keep them at a distance, but not strong enough to pull them back to the ground.
- To reach a point where Earth's gravity is reduced to one-millionth of that on Earth's surface, one would have to be 6.37 million kilometers [ 3.73 million miles] away from Earth (almost 17 times farther away than the Moon).
- Even at that point, you still have not totally escaped the Earth's gravity, it is merely too weak to have much of an effect. In fact, as another NASA site explains, "The effect of gravity extends from each object, indefinitely into space in all directions." The key word here is "indefinitely." No matter how far you go, you can never fully
escape the pull of the Earth, a somewhat comforting thought when faced with the vast and endless expense of space.
- Technically, it goes on forever, BUT it weakens quickly as the distance from the Earth increases. This is one expression of the Inverse Square Law. For gravity, it states that an object that is at a distance twice the radius of the Earth will experience a gravitation pull that is $1 / 4$ the pull at the surface of the Earth, and an object that is 3 r from the Earth will feel a pull that is $1 / 9$ the amount at the surface. So you can see that it drops off pretty fast, but technically it would never be zero, but get asymptotically close to zero. Thus the asteroid will "feel" the gravitational pull of Earth, but at the distance it is, the effect will be VERY small.
- This makes the strength of gravity on the "surface" of the sun (that is, the photosphere, the shiny part we see), 28 times stronger than the force of gravity on the surface of the Earth. Out here, at the distance we orbit the sun, the gravitational pull of the sun is only 0.0006 of the strength of the earth's gravity on the surface of the earth. But that's enough to pull the entire planet around in a big, nearly circular orbit, once per year. And the variation in the strength of the sun's gravitational pull from the part of the earth that faces towards the sun to the part that faces away is partly responsible for the tides of the ocean. The moon's gravity plays a somewhat larger role in the tides. Although it's weaker than the Sun's gravity here, it varies more from one side of the Earth to another.


## Planets - Data Table

|  | Mercury | Venus | Earth | Mars | Jupiter | Saturn | Uranus | Neptune |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| diameter (Earth=1) | 0.382 | 0.949 | 1 | 0.532 | 11.209 | 9.44 | 4.007 | 3.883 |
| diameter (km) | 4,878 | 12,104 | 12,756 | 6,787 | 142,800 | 120,000 | 51,118 | 49,528 |
| mass (Earth=1) | 0.055 | 0.815 | 1 | 0.107 | 318 | 95 | 15 | 17 |
| mean distance from <br> Sun(AU) | 0.39 | 0.72 | 1 | 1.52 | 5.20 | 9.54 | 19.18 | 30.06 |
| orbital period (Earth years) | 0.24 | 0.62 | 1 | 1.88 | 11.86 | 29.46 | 84.01 | 164.8 |
| orbital eccentricity | 0.2056 | 0.0068 | 0.0167 | 0.0934 | 0.0483 | 0.0560 | 0.0461 | 0.0097 |
| mean orbital <br> velocity(km/sec) | 47.89 | 35.03 | 29.79 | 24.13 | 13.06 | 9.64 | 6.81 | 5.43 |
| rotation period (in Earth <br> days) | 58.65 | $-243^{*}$ | 1 | 1.03 | 0.41 | 0.44 | $-0.72^{*}$ | 0.72 |
| inclination of axis <br> (degrees) | 0.0 | 177.4 | 23.45 | 23.98 | 3.08 | 26.73 | 97.92 | 28.8 |
| mean temperature at <br> surface (C) | -180 to | 465 | -89 to | -82 to | -150 | -170 | -200 | -210 |
| gravity at equator (Earth=1) | 0.38 | 0.9 | 1 | 0.38 | 2.64 | 0.93 | 0.89 | 1.12 |
| escape velocity (km/sec) | 4.25 | 10.36 | 11.18 | 5.02 | 59.54 | 35.49 | 21.29 | 23.71 |
| mean density (water=1) | 5.43 | 5.25 | 5.52 | 3.93 | 1.33 | 0.71 | 1.24 | 1.67 |
| number of moons | 0 | 0 | 1 | 2 | 63 | 62 | 27 | 13 |

What is the volume of Earth? (Asked by: Dallas student)

## Answer

The earth is approximately a sphere (actually it is sphere slightly flattened at the poles). Its volume can be calculated if you know its radius. Use the equation for the volume of a sphere which is $V=4 / 3^{\pi} \times$ Radius $^{3}$

The mean radius of the earth is approximately 6.4 million meters (exact $=6.37 \times 10^{6} \mathrm{~m}$ ). Its volume is then:
$(4 / 3) \times 3.14 \times 6400000^{3} \mathrm{~m}^{3}$

This comes to $1,097,509,500,000,000,000,000$ cubic meters. Needless to say, this is very large! Inside of one cubic meter you could fit seven or eight high school students. I know, I teach high school and I have fit eight students in a cubic meter! So, this would be $137,188,690,000,000,000,000$ students. Is your high school this big?


|  | Mercury | Venus | Earth | Msrs | Jupiter | Saturn | Uranus | Neptune | Pluto |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Mean dist. From sun (AU) | 0.39 | 0.72 | 1.00 | 1.52 | 5.20 | 9.53 | 19.18 | 30.08 | 39.34 |
| Inclination of Orbit (deg) | 7.00 | 3.40 | 0.00 | 1.86 | 1.31 | 2.49 | 0.66 | 1.77 | 17.14 |
| Orbital Eccentrically | 0.206 | 0.007 | 0.017 | 0.093 | 0.048 | 0.056 | 0.046 | 0.010 | 0.248 |
| Orbital speed(KM/S) | 47.89 | 35.00 | 29.79 | 24.13 | 13.06 | 9.64 | 6.81 | 5.43 | 4.74 |
| Diameter (KM) | 4.878 | 12.103 | 12.755 | 6.790 | 142.796 | 120.660 | 51.118 | 49.528 | 2.284 |
| Mass ( Earth=1) | 0.06 | 0.81 | 1.00 | 0.11 | 318 | 95 | 14.5 | 17.14 | 0.002 |
| Density(Water=1) | 5.43 | 5.24 | 5.52 | 3.95 | 1.33 | 0.69 | 1.29 | 1.64 | 2.03 |
| Length of Days | 58.7 dys | 243 dys | 23 hrs 56 min | 24hrs 37 min | 9 hrs 56 min | $10 \mathrm{hrs40min}$ | 17 hrs 14 min | 16 hrs 6 min | 6 dys 9 hrs |
| Length of year | 87.97 dys | 224.7dys | 365.26dys | 686.98 dys | 11.86 yrs | 29.46 yrs | 84.07yrs | 164.8 yrs | 248.6 yrs |
| Number of month | 0 | 0 | 1 | 2 | 16 or MORE | $\begin{aligned} & 18 \mathrm{or} \\ & \text { MORE } \end{aligned}$ | 15 or more | 8 or more | 1 |
| Temperature range(0C) | -180 to 430 | 465 | -80 to 60 | -122 to 25 | -150(cluds) | -170(cluds) | -210(cluds) | -110(cluds) | -220 |
| Atmosphere | Almost None | $\mathrm{CO2}$ | Nitrozen Oxygeon | CO2 | Hydrogeon Heliun | Hydrogen Heliun | Hydrogen Heliun Methane | Hydrogen Heliun Methane | None(7) |
| Above : |  |  |  |  |  |  |  |  |  |
| The table above provides a quick way to took up the key facts and statistics of the nine planets of the Solar System. Values given for the inclination and eccentrically are the values for the planet's orbit, multiply the mean distance from the sun value listed in the table by approximately $149,597,900$ kilometers ( $93,000,000$ miles). In order to compute the mass of a planet, multiply the value listed in the t able by approximately $5.9742 \times 10 / 24$ kilograms. The numbers given for length of a day and length of a year are expressed in : Earth" time. |  |  |  |  |  |  |  |  |  |


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| 1eヨ $\times 980$ ६9） <br> OL X カ七を\＆8＇9 |  |  |  | $\begin{gathered} 0 \\ \substack{z_{1} 01 \times 1 \\ Z \varepsilon 80.1} \end{gathered}$ |  |  |  |
| $\begin{aligned} & \text { دеヨ } \times 6086 \cdot \varepsilon) \\ & \text { Z98'gz } \end{aligned}$ |  | $\begin{gathered} \text { (丩นеヨ } \times \varepsilon \varepsilon \angle 6^{\circ} \text { OL) } \\ 116^{\prime} 69 \end{gathered}$ |  | $\begin{gathered} 0 \\ 0 \\ 0 \quad 1 \angle \varepsilon^{\prime} 9 \end{gathered}$ | $\begin{gathered} (4 \text { यe } \exists \times 66+6.0) \\ 8.590^{\circ} 9 \end{gathered}$ |  | （wy）snipey <br>  ue？ |
|  | ZZヤ＇999＇9Zヤ＇し | เマ8＇0ヤを＇8LL | †て8＇દャ6＇Lて乙 | $\begin{gathered} z 9 z^{\prime} 8 \\ 69^{\prime} 6 \downarrow \downarrow \end{gathered}$ | 9Lヤ＇60Z＇801 | LZて＇606＇L9 | （3iquo jo sixe 1о［еш！ues） （ux）uns euz шо．！әэuets！a |
| （иәлеәН рор уэәәд пие $\wedge \overline{\text { snued }}$ |  јо рор ие шоу）（！urчS）$\overline{\text { unłes }}$ |  <br>  <br>  |  |  |  | （әu®］ <br>  <br>  |  |

STIVLAG－SLANNTId－WGLSAS \＆VTOS

| Ecliptic |  |  |  |  |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Inclination of Equator to Orbit | $0^{\circ}$ | $177.3^{\circ}$ (retrograde rotation) | $\begin{gathered} 23.439 \\ 3^{\circ} \end{gathered}$ | 25.2 | $3.1^{\circ}$ |  | $26.7^{\circ}$ | $97.8^{\circ}$ (retrogre rotation) |
| Minimum/Maxi <br> mum <br> Surface <br> Temperature | -173/427 | 462 | -88/58 (min/m ax) | -87 to -5 |  |  |  |  |
| Major Atmospheric Constituents |  | Carbon Dioxide, Nitrogen | $\begin{aligned} & \text { Nitroge } \\ & \text { n, } \\ & \text { Oxygen } \end{aligned}$ | Carbon Dioxide, Nitrogen, Argon | Hydrogen, Helium |  | Hydrogen, Helium | Hydrogen, Heli Methane |
| Moons | None | None | 1 moon | 2 moons | 62 moons |  | 62 moons | 27 moons |
| Rings | No | No | No Mn గ్రహములు: వాటి వివరాలు |  | గ్రహములు: వాటి వివరాలు |  |  | $/$ |
| Primary Source: Cox, Arthur, ed., Alleı Secondary Source: JPL Solar System |  |  | గ్రహం పేరు | వ్యానం <br> (รి. మిలeో) | దినం (భూమి <br> రోజులు,గంటలు) | నంవత్సరం (భూ\%ంవత్సరం) | సూర్యుడు నుండి దూoం కి. మీ |  |
| Last Updated: 11 May 2011 |  |  | బుధుడు | 4880 | 58.6 రోజులు | 88 / రోజులు | 5,79,00,000 |  |
|  |  |  | శుక్రుడు | 12109 | 243 రోజులు | 224 రోజులు | 10,82,30,000 |  |
|  |  |  | భూమి | 12835 | 24 గంటలు | 365 రోజులు | 14,9590,000 |  |
|  |  |  | కుజుడు | 6780 | 24.6 గంటలు | 687 రోజులు | 2277,20,000 |  |
|  |  |  | గురుడు | 142,725 | 9.8 గంటలు | 11.6 సంవత్సరాలు | 77,81,20,000 |  |
|  |  |  | శని | 120050 | 10.2 గంటలు | 29.4 సంవత్సరాలు | 142,83,09000 |  |
|  |  |  | యuరేనస్ | 51705 | 17.30 గంటలు | 84 సంవత్నరాలు | 287,27,00,000 |  |
|  |  |  | నెహ్వూన్ | 49520 | 15.8 గంటలు | 11 సంవత్సరాలు | 449,81,00,000 |  |
|  |  |  | ప్లూటో | 2215 | 6.3 రోజులు | 247.8 సంవత్సరాలు | 591,43,00,000 |  |

67000 మైళ్ళు / గంటకు 18.5 మైళ్ళు / సెకనుకు
1/81.5 బరువు
190000000 (పందొమ్మిది కోట్లు) ఈ సంఖ్య క్రింద ఇంకా 20 సున్నాలు
పెడితే వచ్చే సంఖ్య టన్నులు
5870 ప్రక్కన 19 సున్నాలు పెడితే వచ్చే సంఖ్య టన్నులు

భూమి ఆకర్షణ శక్తి భూమి నుండి

భూమి నడక వేగం
భూమిలో చంద్రుడు
సూర్యుని బరువు
భూమి బరువు
భూమి బరువు

Primary Source: Cox, Arthur, ed., Alle,
Secondary Source: JPL Solar System
Last Updated: 11 May 2011

